

Status of Euclid and perspective of CMB cross-correlations

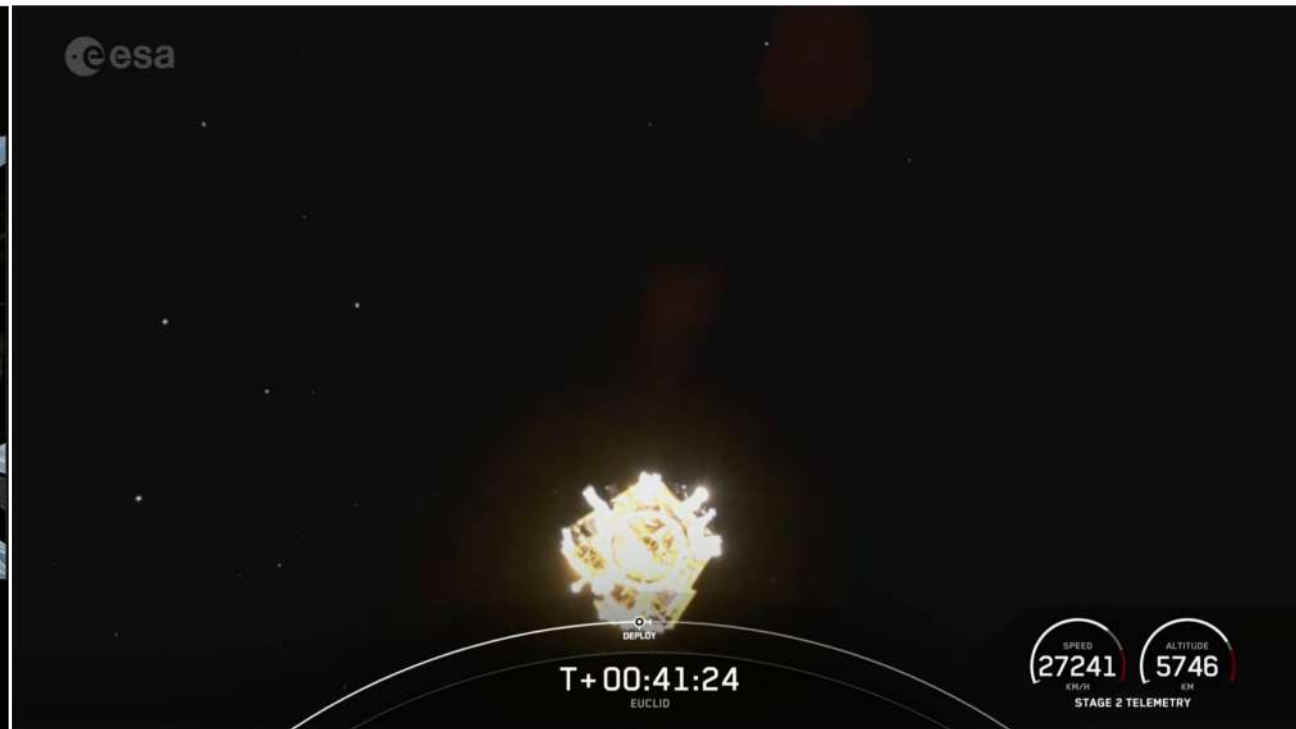
Workshop “CMB Day 2”
ASI - 17 October 2023

Marina Migliaccio



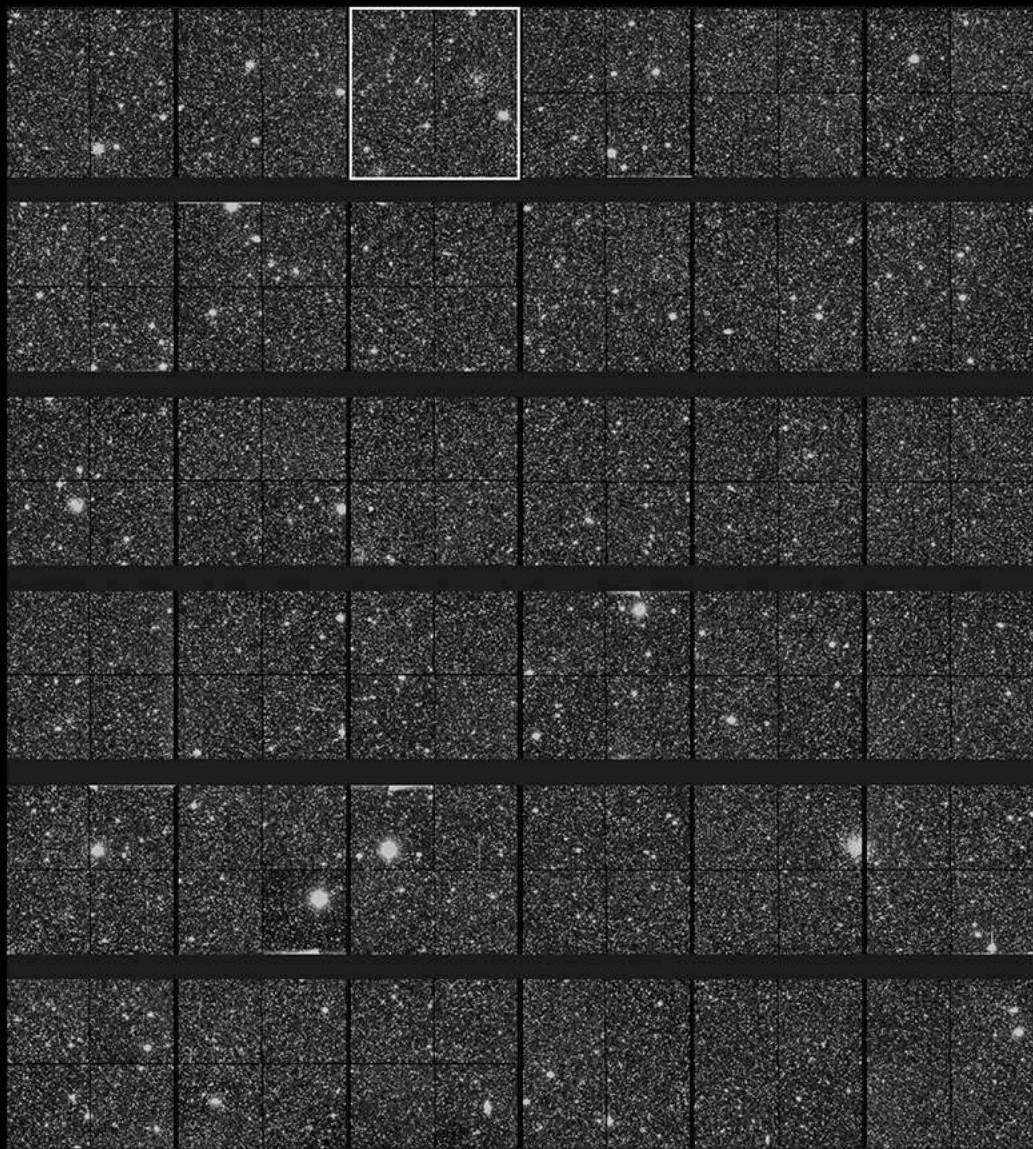
Successfully launched on a SpaceX Falcon 9 rocket from the Cape Canaveral Space Force Station in Florida.
On July 1st, 2023 at 5.12 PM CEST



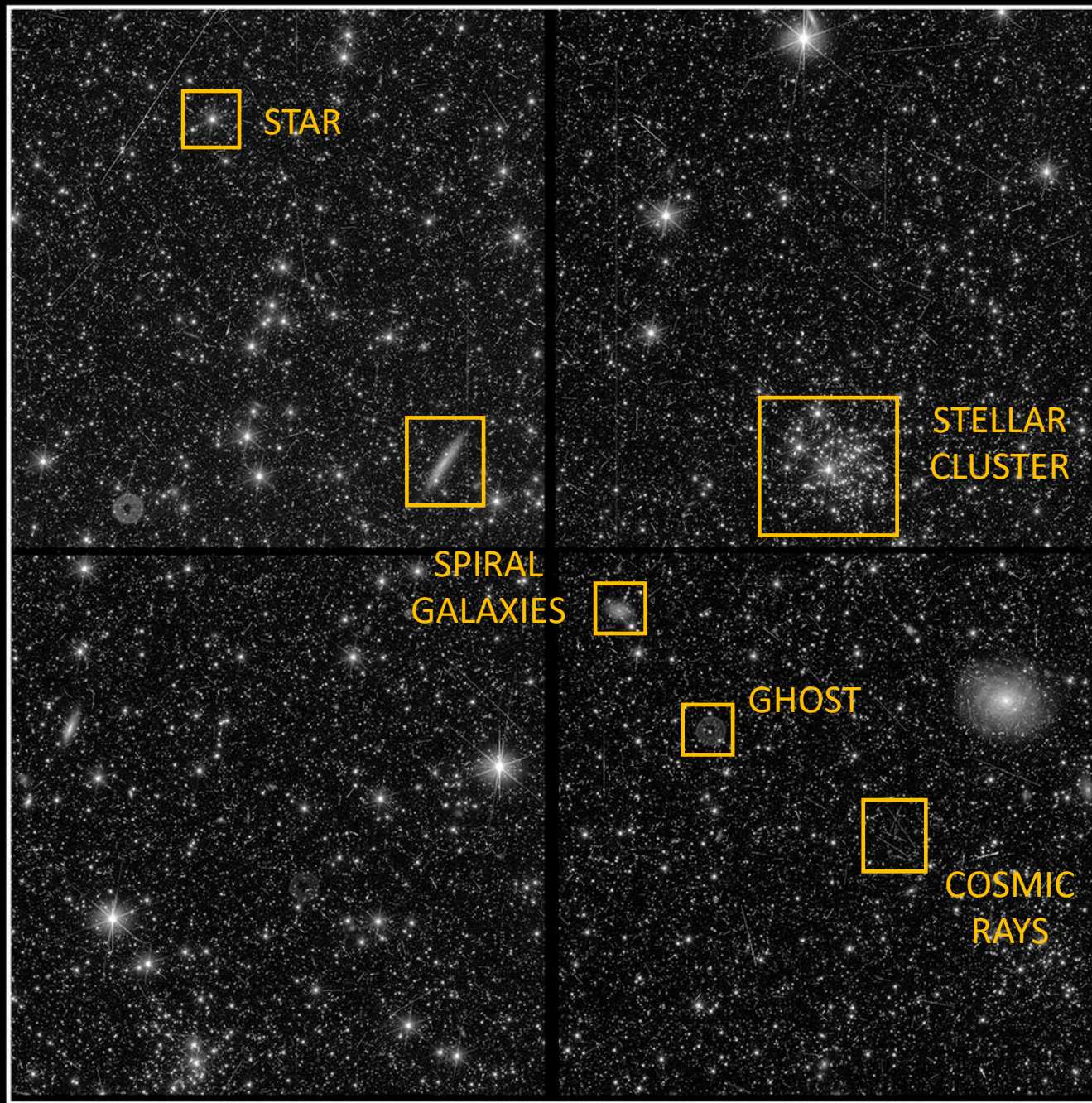
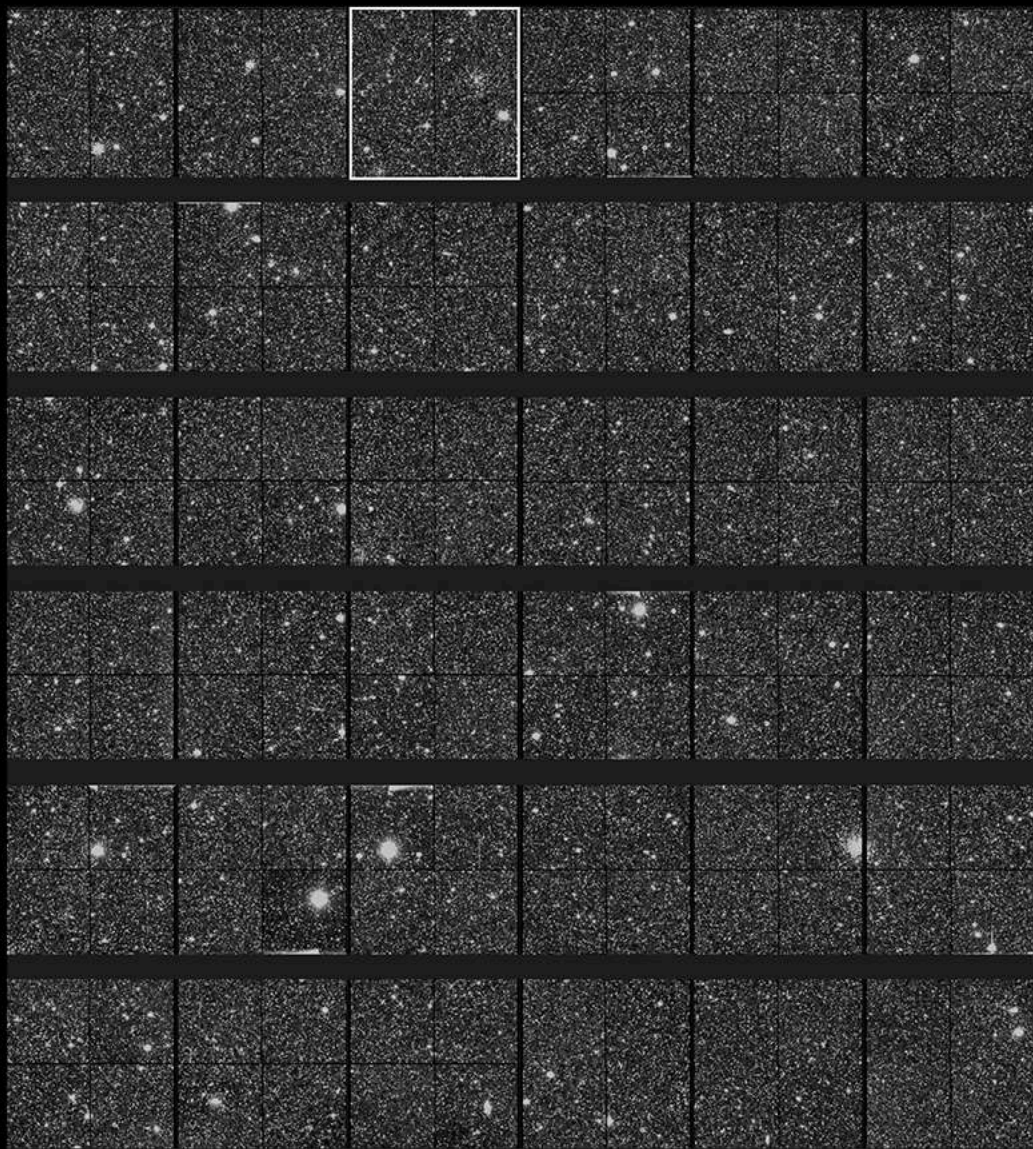


- **5 October:** End of commissioning, updated guiding software and PV resumes.
- **19 August:** Commissioning extended, troubleshooting the fine guidance sensor
- **12 August:** Performance Verification starts
- **31 July:** First images released
- **28 July:** Euclid focused and at destination orbit
- **15 – 18 July:** VIS and NISP see first light
- **2 - 4 July:** On the way to L2 - Commissioning starts

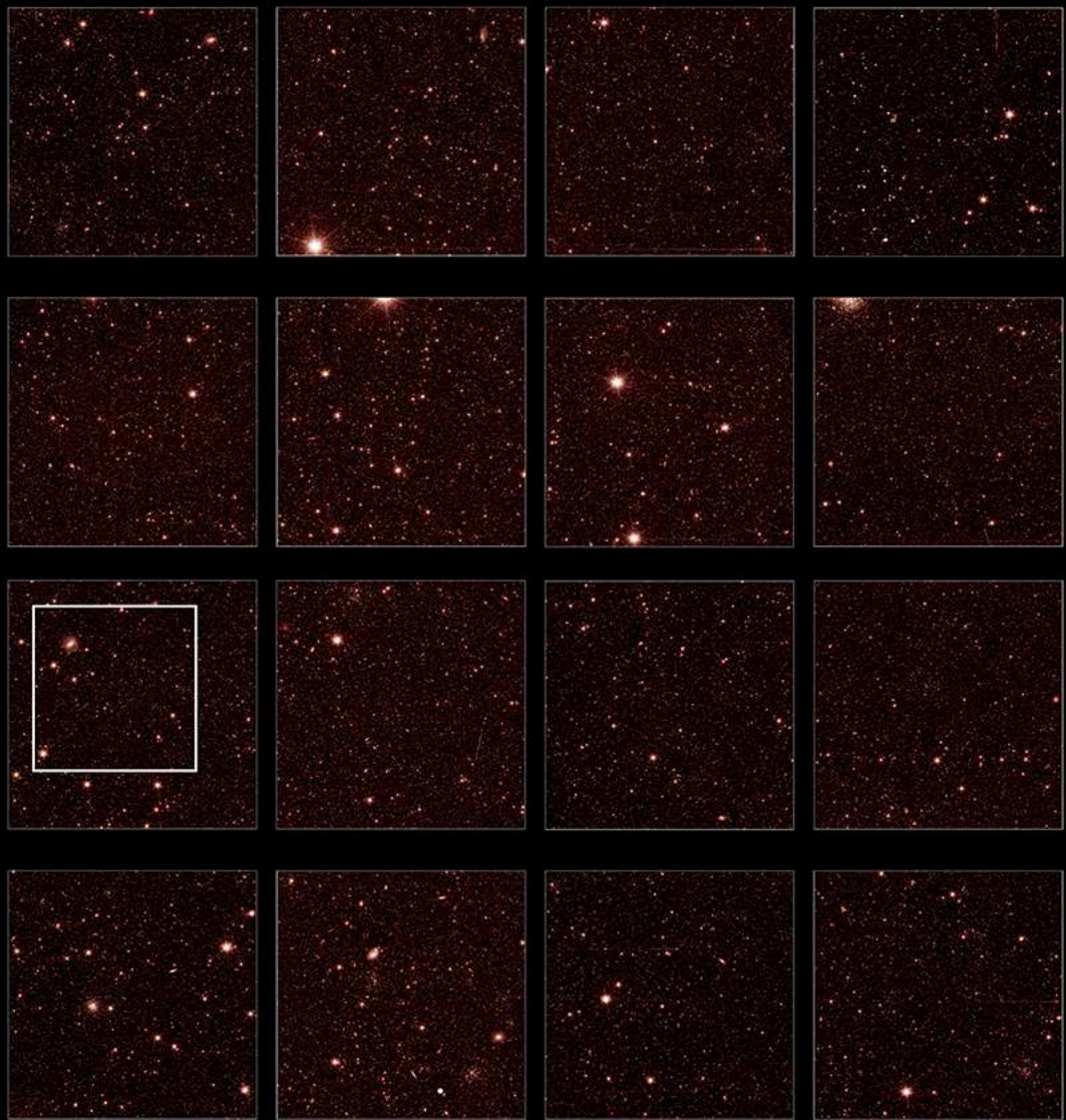
EARLY COMMISSIONING TEST IMAGE, VIS INSTRUMENT



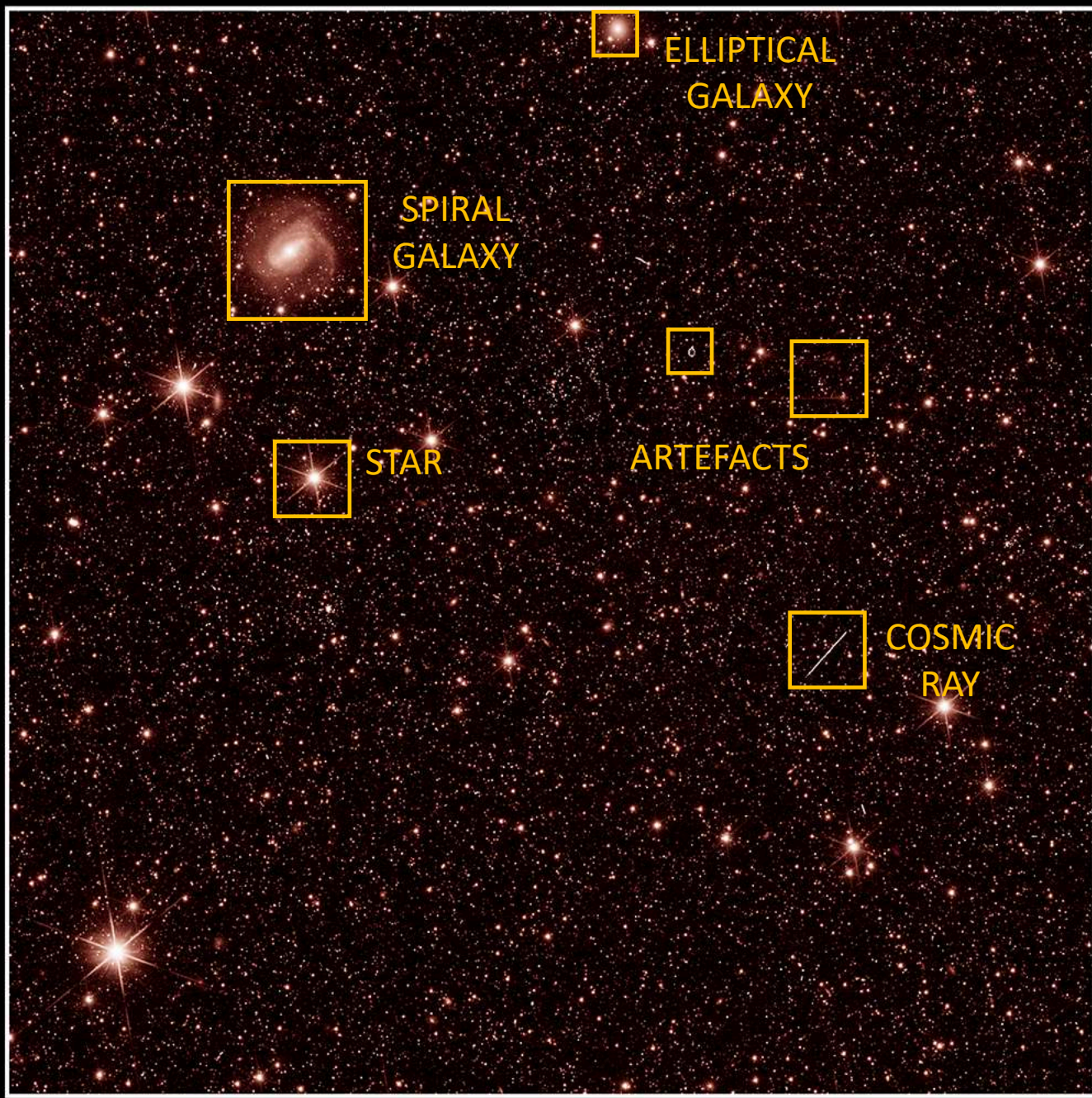
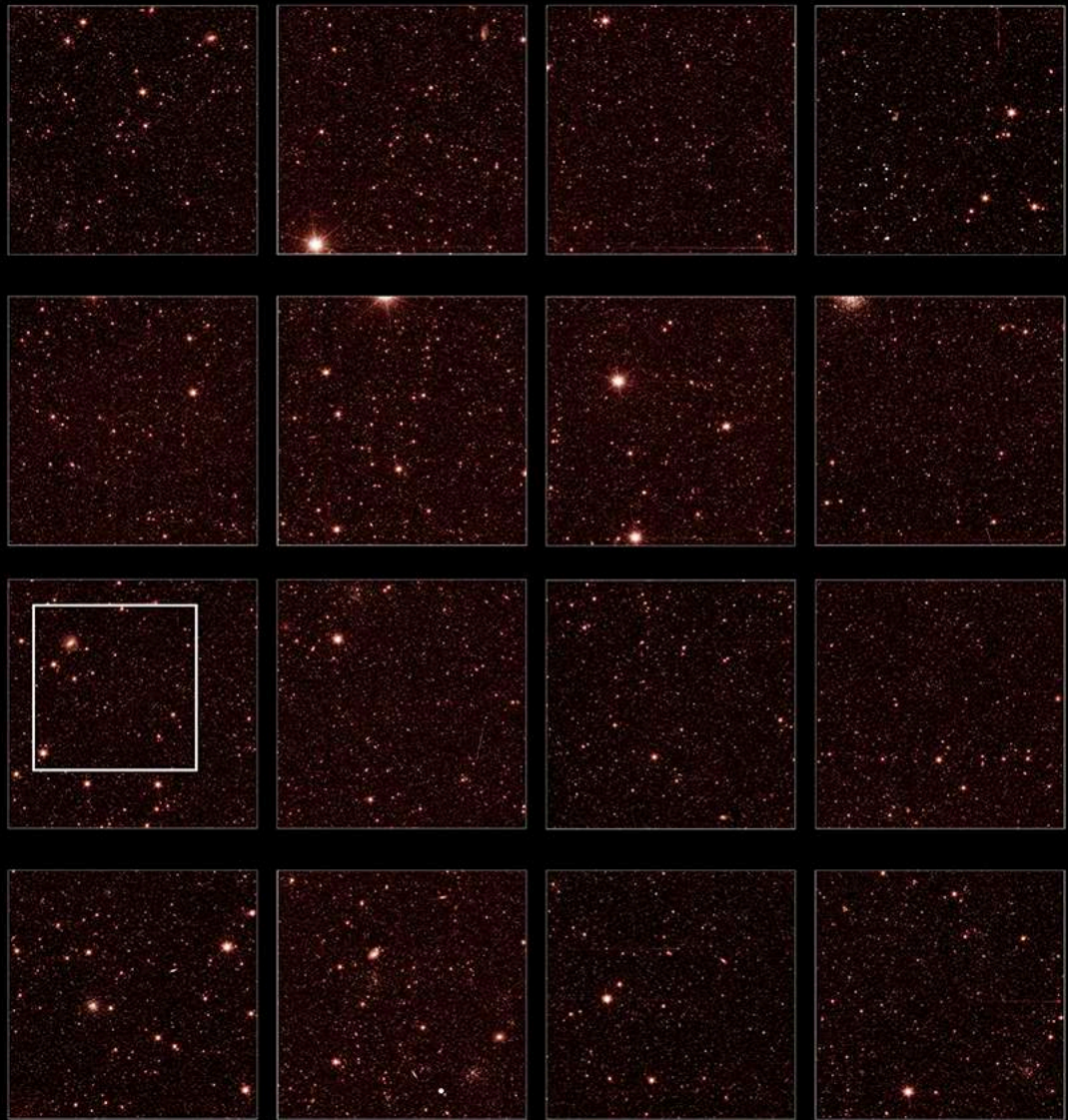
EARLY COMMISSIONING TEST IMAGE, VIS INSTRUMENT

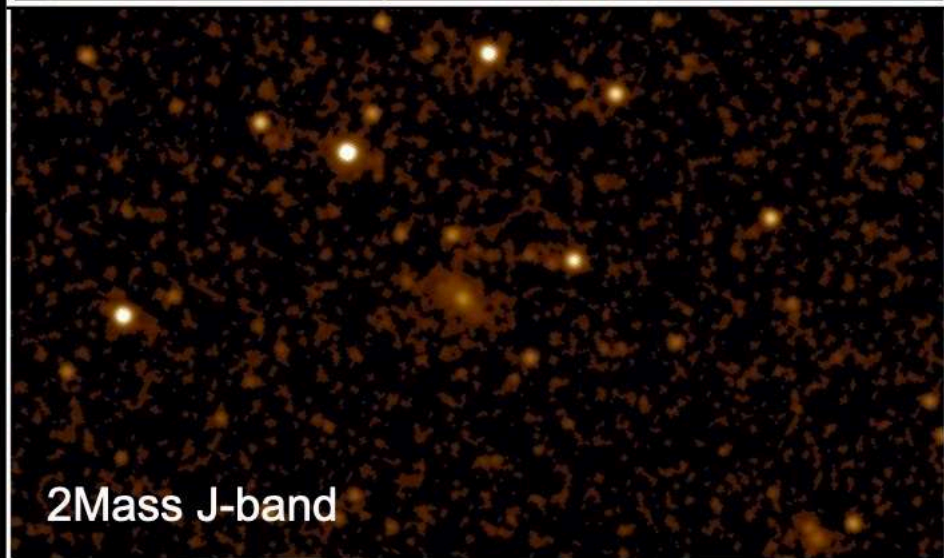


EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT



EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT







- Medium Class ESA Mission
- Visible to near-infrared (with spectroscopy) space telescope operating in L2
- Image billions of galaxies out to a distance of 10 billion light years ($z \sim 2$) across 36% of the sky
- Primary Probes: **Galaxy Clustering** and **Cosmic Shear** to trace the expansion history and the growth rate of cosmic structures with unprecedented precision
- Other probes: **Clusters**, **Voids**, **CMBX**, **Strong Lensing**, and more

ESA science requirements: Tight constraints on **deviations from Λ CDM**, **neutrino masses**, **initial conditions**.
[Euclid red book \(2011\) arXiv: 1110.3193](#)

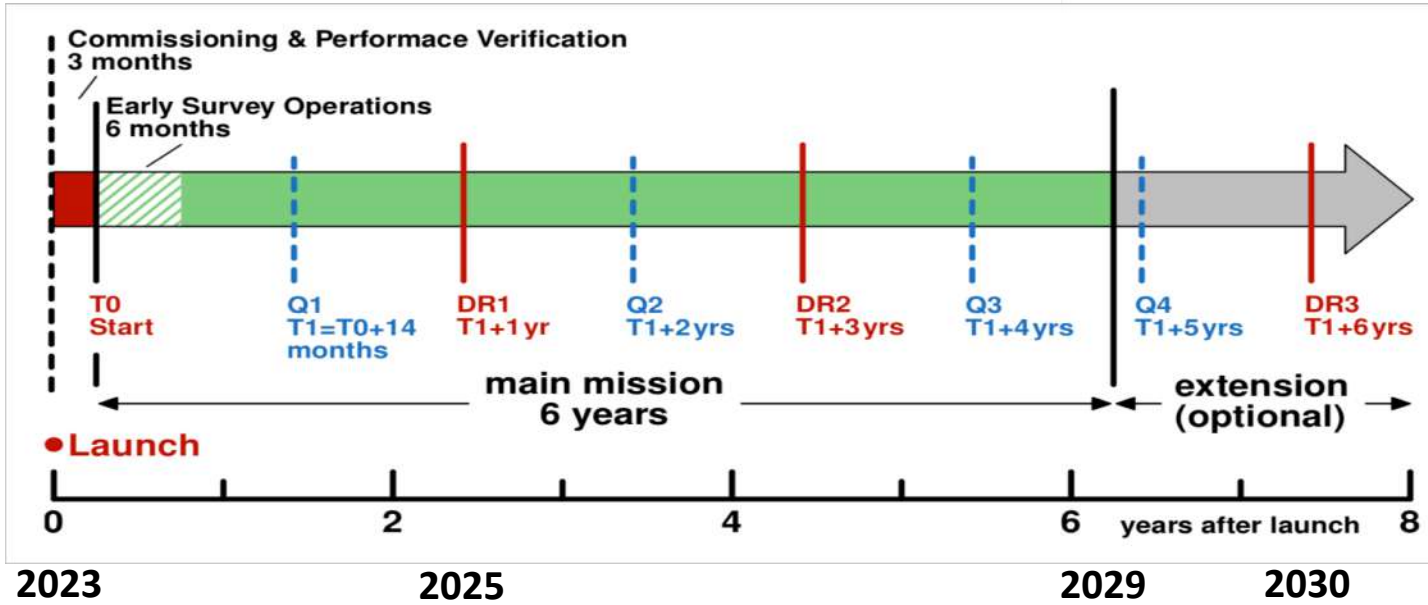
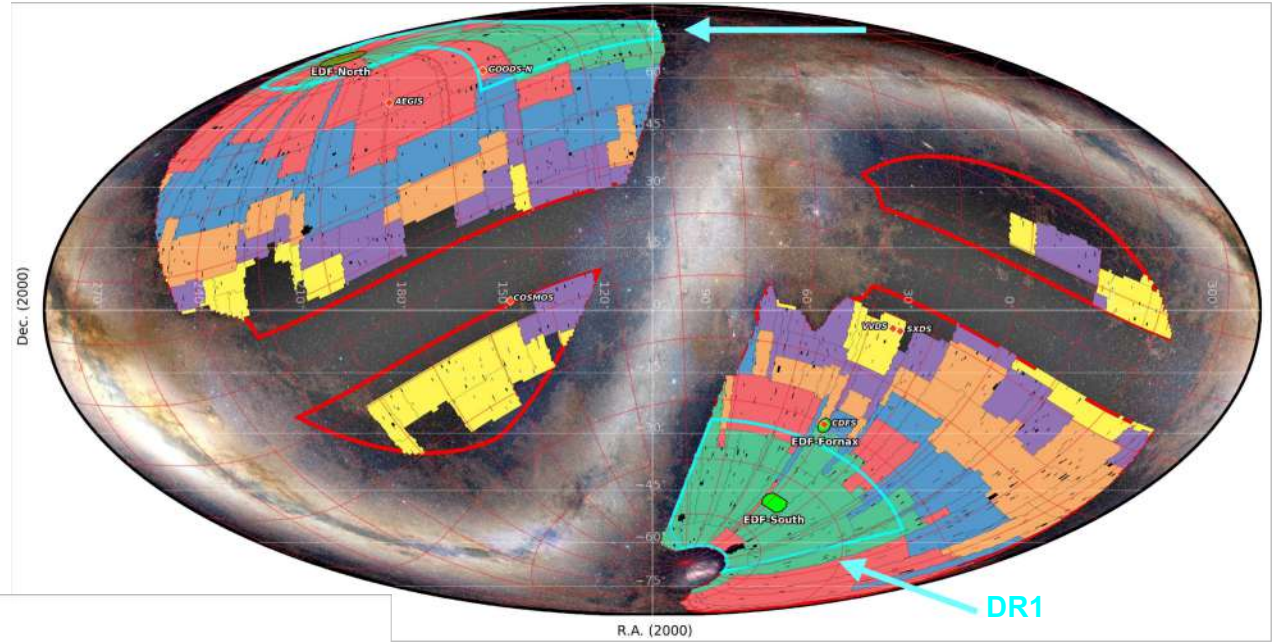
	Modified Gravity	Dark Matter	Initial Conditions	Dark Energy		
Parameter	γ	m_ν/eV	f_{NL}	w_p	w_a	FoM
Euclid Primary	0.010	0.027	5.5	0.015	0.150	430
Euclid All	0.009	0.020	2.0	0.013	0.048	1540
Euclid+Planck	0.007	0.019	2.0	0.007	0.035	4020
Current	0.200	0.580	100	0.100	1.500	~ 10
Improvement Factor	30	30	50	>10	>50	>300

BUT will do much more..

More than 30 Euclid Preparation Papers already published and many more soon to come



Timeline of Euclid

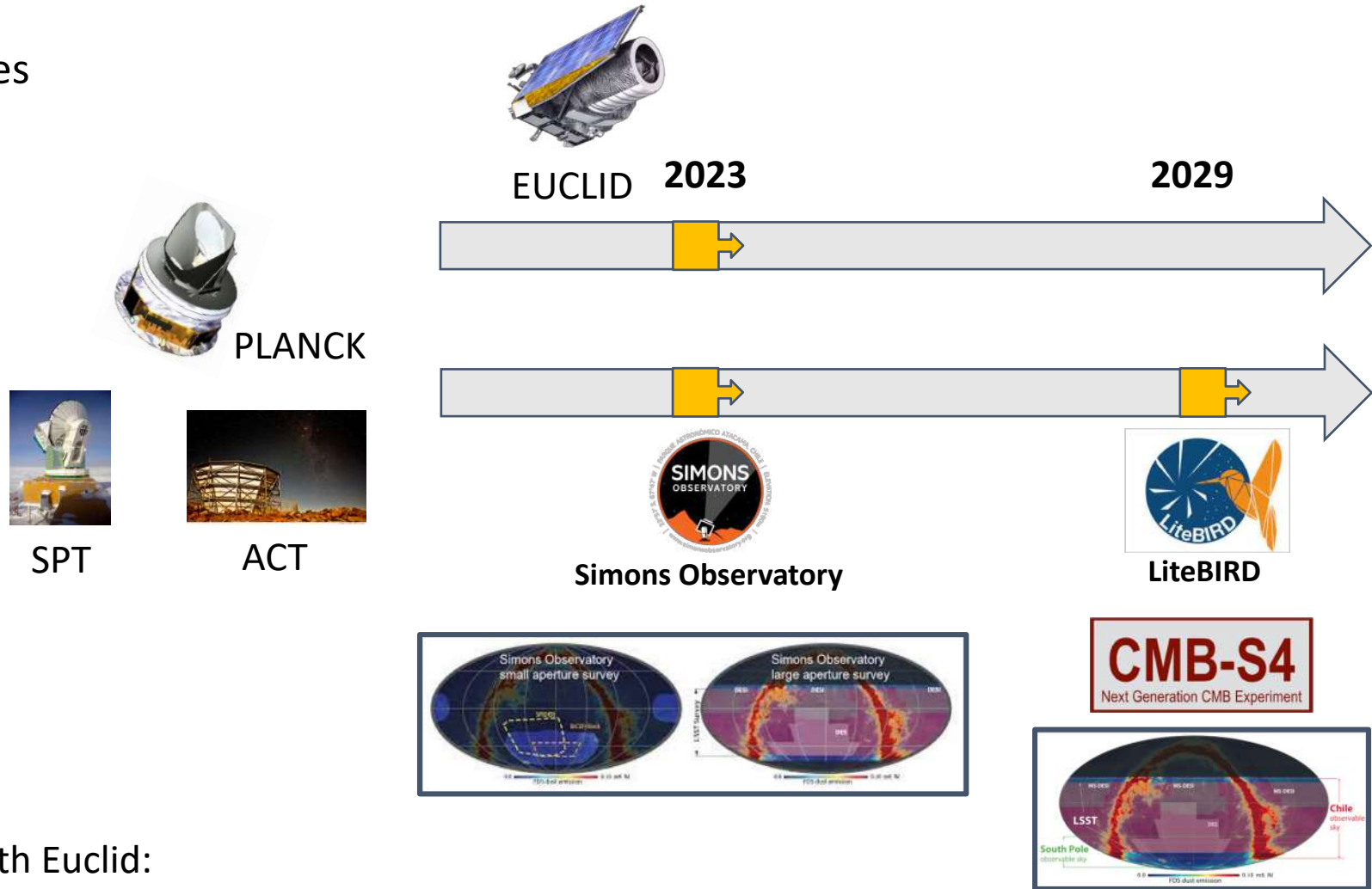


Euclid Wide Survey

Timeline of CMB Probes

Upcoming CMB experiments will be targeting:

- Polarization
- Secondary Anisotropies



For the cross-correlation with Euclid:
maps of the northern sky only from space experiments!



Euclid and CMB joint analyses

COMBINATION and CORRELATION of the upcoming Euclid data with CMB measurements will provide the **largest lever arm of epochs**, ranging from recombination to structure formation and the late-time accelerated expansion of the Universe.

- **Maximise the return of information** in constraining the **cosmological parameters**.

$$\{C_{\ell}^{TT}, C_{\ell}^{TE}, C_{\ell}^{EE}\} + \{C_{\ell}^{TG}, C_{\ell}^{\kappa\kappa}, C_{\ell}^{\kappa G}, C_{\ell}^{\kappa\gamma}, C_{\ell}^{GG}, C_{\ell}^{\gamma\gamma}, C_{\ell}^{\gamma G}\}$$

- **Cross-Correlations** originate from **CMB secondary anisotropies**
 - photons interact with structures **gravitationally** (Lensing, late-time ISW effect) or by **scattering** (SZ effect).
- Cross-correlation science is transitioning from detection regime to powerful cosmological probe.
 - **Break degeneracies** between cosmological and astrophysical parameters.
 - **Control uncorrected systematic effects** in the data.
 - Build **novel estimators** for cosmology and astrophysics.



CMBX Science Working Group

Leads: C. Baccigalupi, G. Fabbian, S. Ilic (deputy)

Deep involvement of the Italian CMB community

with leadership roles in several Key Projects

Activities

Aosta: CMB-N-Body & Ray Tracing, Simulations, Analysis

Bologna: CMB-N-Body Simulations, Analysis, Estimators & Likelihood, Theory

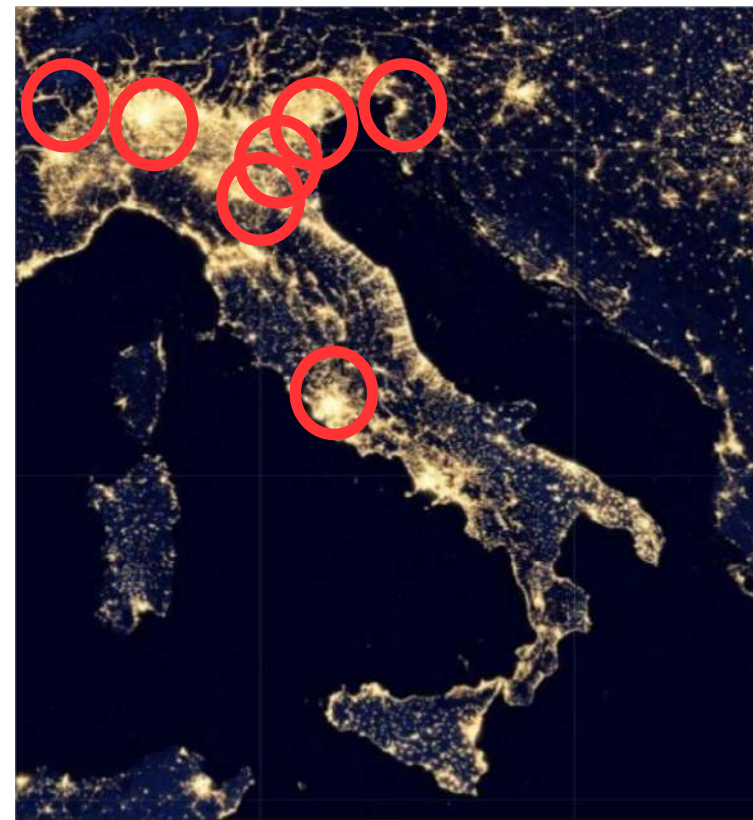
Ferrara: Estimators & Likelihood

Milano: CMB-N-Body & Ray Tracing Simulations, Analysis

Padova: Estimators, Analysis, Theory

Rome: Estimators & Likelihood, Simulations, Analysis

Trieste: Analysis of CMB-N-Body Simulations



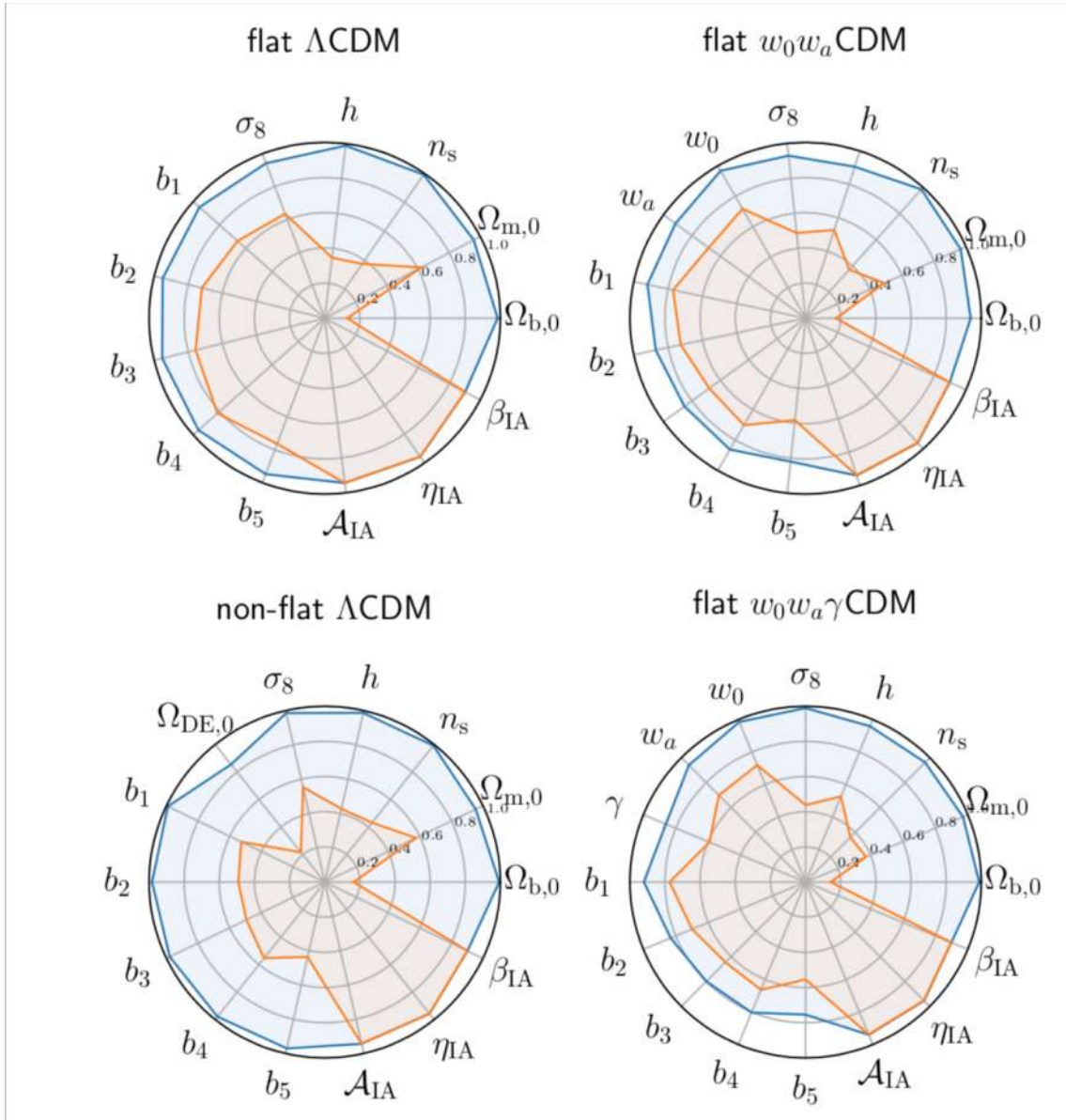
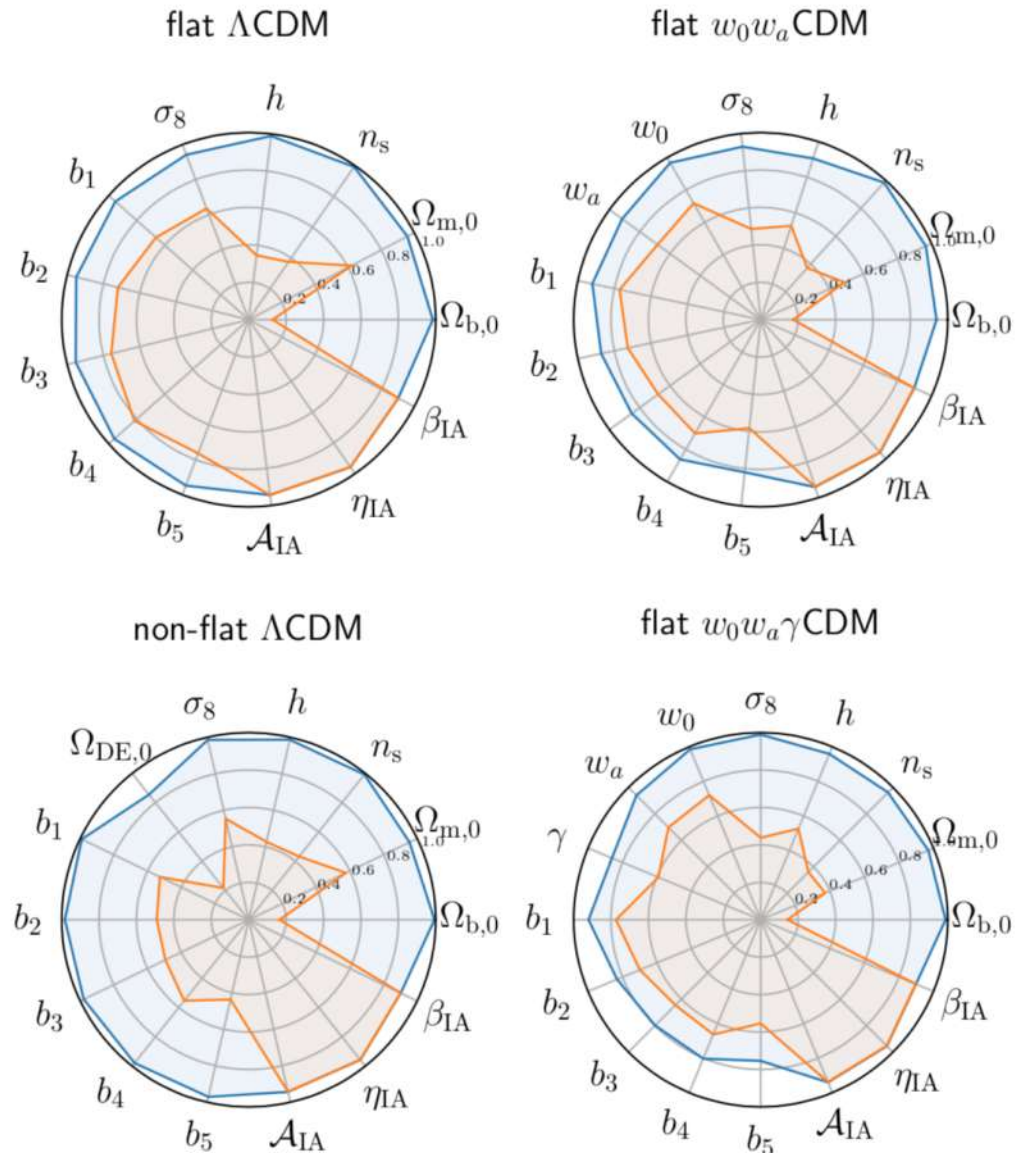


Fig. 6. Ratio of predicted 1σ uncertainties (see end of Sect. 5.2) showing how constraints are tightened after adding CMB lensing (blue) or all CMB probes (orange) when compared to the *Euclid*-only constraints (black outer rim), assuming a pessimistic *Euclid* scenario and SO-like CMB data, for four selected cosmological models (*from top to bottom, left to right*: flat Λ CDM; flat $w_0 w_a$ CDM; non-flat Λ CDM; and flat $w_0 w_a \gamma$ CDM).



**A new Key Project paper in preparation
on CMBX Extended Forecasts**

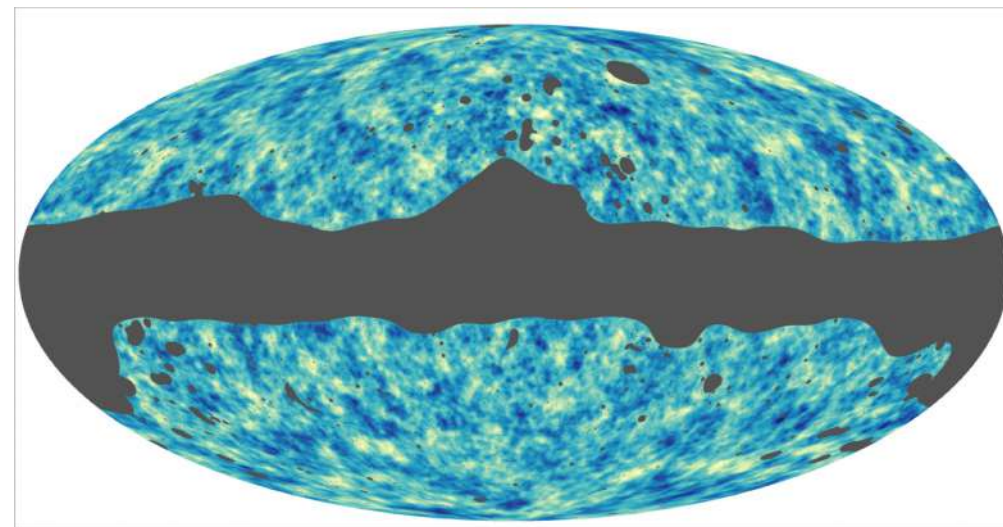
- **Non-standard cosmologies**
- **Characterization of Euclid Probes**

Fig. 6. Ratio of predicted 1σ uncertainties (see end of Sect. 5.2) showing how constraints are tightened after adding CMB lensing (blue) or all CMB probes (orange) when compared to the *Euclid*-only constraints (black outer rim), assuming a pessimistic *Euclid* scenario and SO-like CMB data, for four selected cosmological models (from top to bottom, left to right: flat Λ CDM; flat $w_0 w_a$ CDM; non-flat Λ CDM; and flat $w_0 w_a \gamma$ CDM).

Euclid and CMB cross-correlation analyses

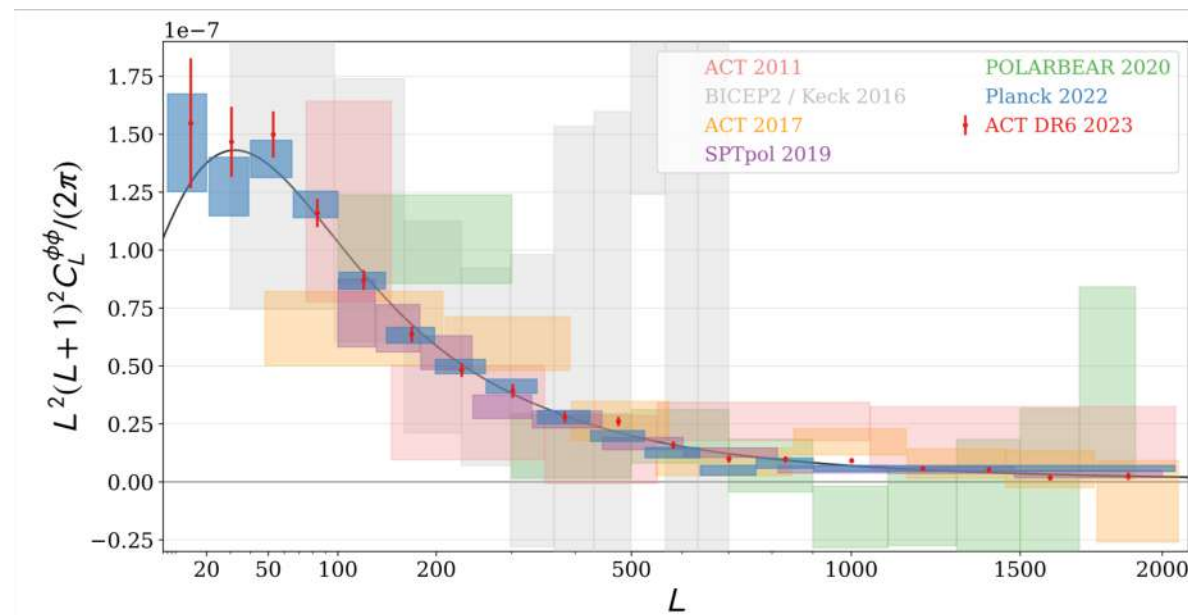
CMB Lensing

- Unbiased tracer of the mass distribution in the Universe integrated along the line of sight.
- In the future complementary reconstructions from temperature and polarization CMB data.

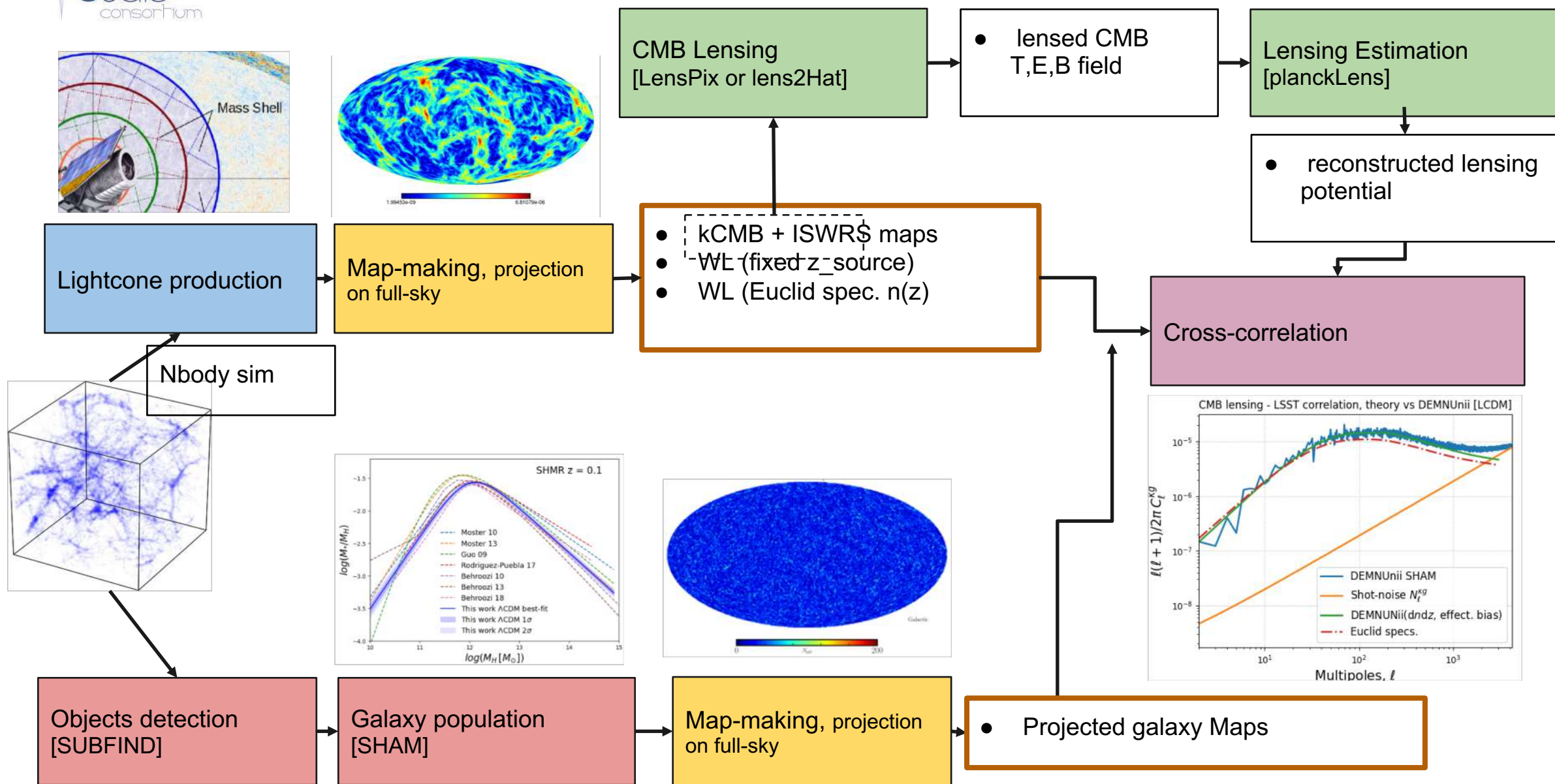


CMB Lensing x Galaxy Clustering

- now **detected from radio frequencies to γ -rays**.
- κG can be measured with $S/N > 100$ with Euclid x SO
- powerful **confirmation of GR** on cosmological scales
- breaks degeneracy with **galaxy bias**
- constraints on **late-time expansion** and **clustering** of matter



CMBX-sims End-to-end Pipeline (left to right)



Euclid and CMB cross-correlation analyses

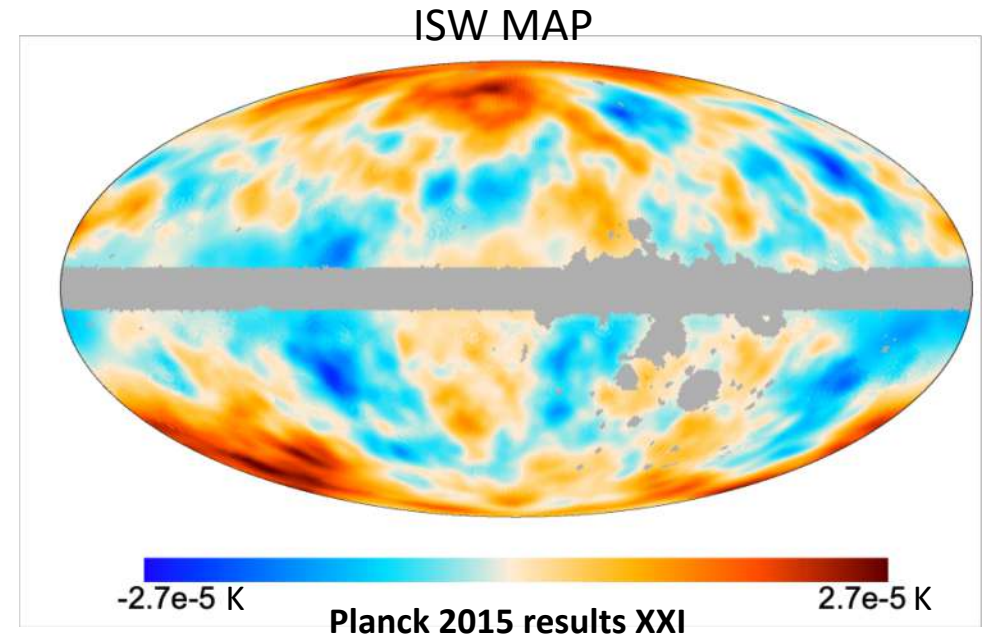
Late-time Integrated Sachs-Wolfe Effect

Time-varying gravitational potentials along the photon path induce CMB temperature anisotropies.

→ the ISW effect can only be measured by cross-correlating CMB temperature anisotropies with LSS tracers

TG (ISW) can be measured with $S/N \simeq 4$ with Planck reaching $S/N \simeq 4.6$ with LiteBIRD including polarization data

- Dark energy parameters
- Modified gravity models
- Primordial Non-Gaussianity



Optimal extraction of the iSW signal with Euclid:

- large number of galaxies
- wide coverage in sky area and redshift
- high-degree of correlation
- tomographic information

Euclid and CMB cross-correlation analyses

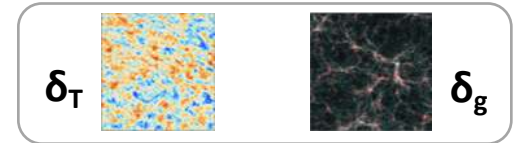
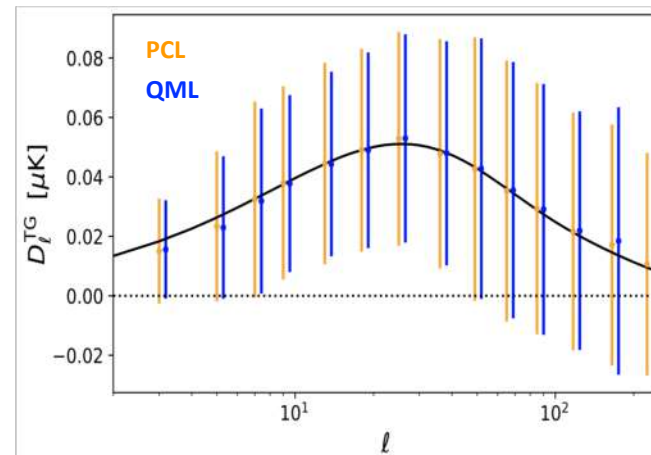
Late-time Integrated Sachs-Wolfe Effect

Key Project

Development of an end-to-end analysis pipeline

- Theoretical predictions
- Simulations
- Estimators
- Likelihood
- Cosmological parameters

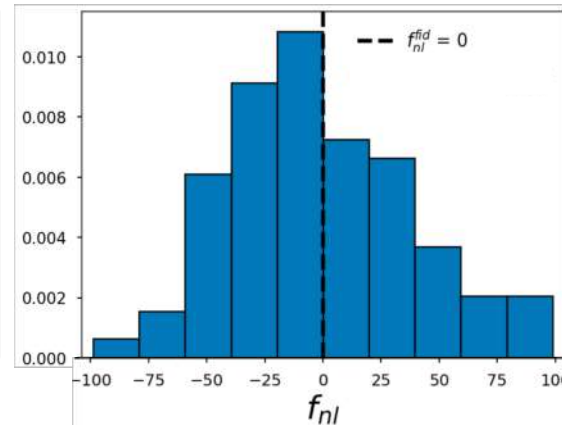
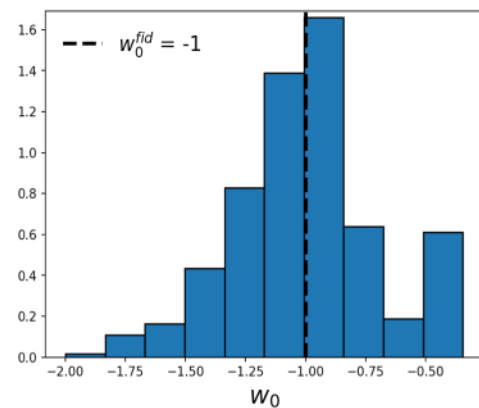
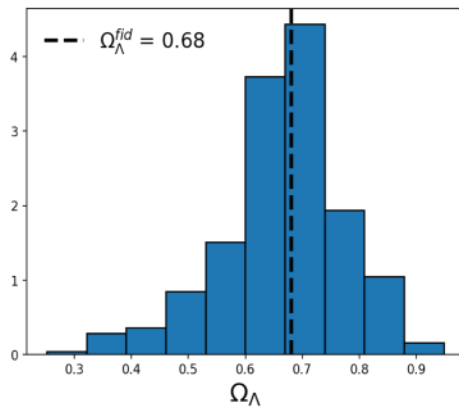
+ Needlet based estimator
Cross-correlation functions



C^{Tg}

Likelihood
 $P(\text{model} | \text{data}) \propto P(\text{data} | \text{model}) P(\text{model})$

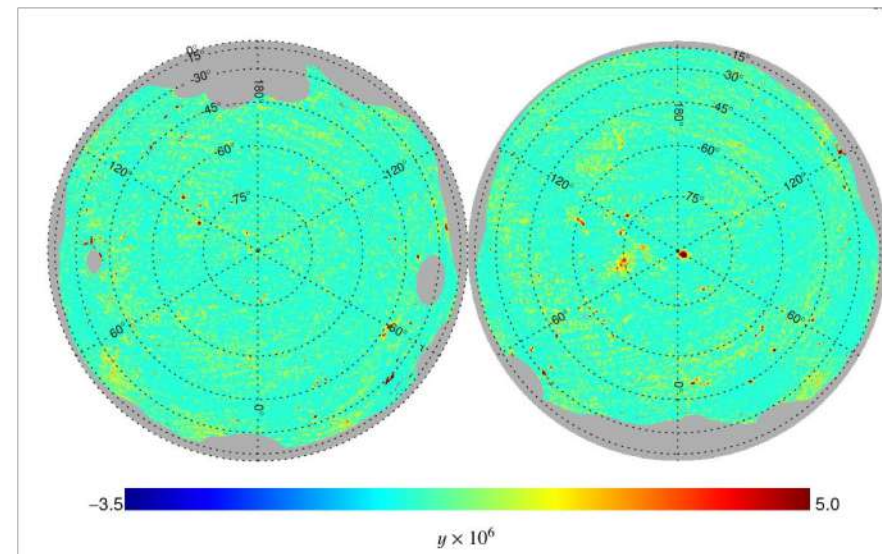
$\Omega_\Lambda, w_0, w_a, \gamma, \dots$



Euclid and CMB cross-correlation analyses

Thermal and Kinetic Sunyaev Zel'Dovich Effect

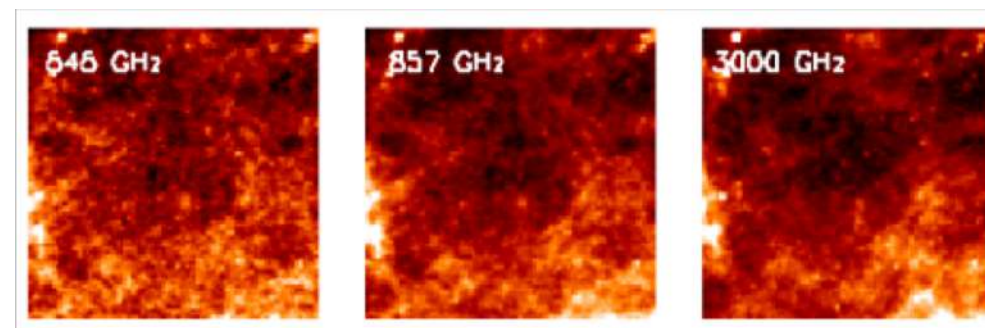
- Mapping gas profiles
- Hot gas tomography



Planck 2015

Cosmic Infrared Background

- Star formation at high redshift



Planck 2013



Euclid is flying!

Will probe the dark sector with unprecedented accuracy.

Combination of the primary probes is crucial to reach the expected scientific targets.

Synergy with external probes will be key to fully exploit the dataset to come.

CMB CAN PLAY AN IMPORTANT ROLE: NEED TO BE READY.

With ideas, theoretical developments, numerical simulations, data analysis pipelines, structured support of the activities, MoU with CMB experiments, ...